## Introduction to the Standard Model

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#### What are we made of?



#### **Elementary Particles**

- They are the basic constituents of matter
- They cannot be subdivided (they have no internal structure)
- Everything we see and touch is made of only three particles: up quark, down quark, and the electron



#### The Forces of Nature

• Why are atoms this way?



- Electrons orbit the nucleus thanks to the electric force
  - Question #1 : What is a force?

It's an exchange of particles (in this case photons)

#### Analogy

- Two persons standing on an ice skating ring
- How can I move the other person without touching her?

**Answer: by throwing something!** 



Note: to understand an attractive force you would have to imagine a ball with negative energy

#### The Forces of Nature

• Why are atoms like this?



- Electrons orbit the nucleus thanks to the electric force
  - Question #2: how come the protons in the nucleus do not repel each other?

There is another force (only between quarks) that is stronger than the electric force: the strong force

#### The Strong Force

• As the name suggests, the strong force is very strong!



proton (positive) neutron (neutral)

- Occurs only between particles that have "color" (quarks)
- Binds protons and neutrons individually and in a group inside the nucleus
- Is transmitted through the exchange of **gluons**

- So far we've only talked about 5 particles
- But that's just the tip of the iceberg

mass → ≈2.3 MeV/c<sup>2</sup> charge → 2/3 U g spin  $\rightarrow$  1/2 gluon up ≈4.8 MeV/c<sup>2</sup> DUARKS -1/3 d 1/2 photon down 0.511 MeV/c<sup>2</sup> Mediator of the -1 e electromagnetic 1/2 force electron Constituents of "regular" matter

Mediator of the

strong force

• There are at least 17 fundamental particles!



Nota: the graviton is not listed

• Even though we tend to think of these particles as little marbles, they behave in a much stranger way



- They obey Quantum Mechanics
  - They can be in multiple places at once
- Many of them are unstable





- They decay in a few microseconds or less!
- This is why we only find two types of quarks (up, down) and one type of lepton (electron) in conventional matter
- But we can produce them in the lab and study them before they disintegrate

• Each particle has its own antiparticle (identical in every way, but with opposite charge):



When a particle meets its antiparticle, they annihilate each other



 Our universe contains mostly matter, but we can produce antimatter in the lab

#### How are new particles created?

 The only way to study unstable particles is by creating them in the lab. How?



# (c is the speed of light in vacuum)

- A little bit of mass can be converted into A LOT of energy \_\_\_\_\_
- It takes a lot of energy to create A LITTLE BIT of mass



#### How are new particles created?

- Need to concentrate a big amount of energy
- The best way to do that currently is with the Large Hadron Collider



#### How are new particles created?

 LHC: two proton beams, going around in opposite directions, in a 27 km circumference ring that is 100 m underground



The most complex machine ever built!

Why so big? Magnets can only be made so strong...

Complex detectors surround each interaction point

#### Example: the ATLAS detector



#### Invisible particles?

- The detectors at the LHC can be used to study almost all the particles of the Standard Model
  - Even the Higgs, whose discovery was announced in 2012!



• The one notable exception? Neutrinos....

#### The Most Abundant Particles

- We mentioned that all conventional matter is made of two types of quarks (up, down) and electrons
- You would think that these are the most abundant particles in the universe



For every electron and quark in the universe there are

# 10 000 000 000

neutrinos

If we want to understand our universe, we need to understand neutrinos

#### Our Universe is Full of Neutrinos!

- Our universe contains an extremely large amount of neutrinos
  - Every cm<sup>3</sup> contains about **300** neutrinos that are relics of the big bang (not yet detected)



Our bodies produce hundreds of millions of neutrinos per day
 v



#### Our Universe is Full of Neutrinos!

 The Sun produces so many neutrinos that every second we are traversed by approximately **100 000 000 000 000** neutrinos (one hundred million million)



#### Our Universe is Full of Neutrinos!

#### • Am I going to die?

#### • Yes, but not because of neutrinos...

of 100 000 000 000 000 000 solar neutrinos that go through our bodies, the number of them that interact (that exchange a particle with a proton, neutron or electron in our body) is:



The rest just breeze through and do **absolutely nothing** (literally)

(of course, if these were protons, it would be a different story)

- In fact, a neutrino can go through a light-year of lead!
- Why do neutrinos interact so weakly with other particles?



• Neutrinos do not interact via the strong force



• Neither do they interact via the electromagnetic force (they have no charge)



 They only interact via the weak force, which is mediated by the W and Z bosons



#### Flashback

• Going back to the previous diagram:



#### Why is the weak force so weak?

- What causes the weak force to be so weak?
  - We saw that a force is an exchange of particles
  - The greatest difference between the weak force and the others is that its mediators (W, Z) are extremely heavy

If interacting via the electromagnetic force can be compared to exchanging little marbles, then interacting via the weak force would be akin to exchanging this:

The W and Z bosons are approximately 100 times heavier than a proton



#### Why is the weak force so weak?

 A good resource to understand this: "Why is the weak force weak" in youtube, with Dr. Don Lincoln

![](_page_29_Figure_2.jpeg)

#### The Weak Interaction

- If the weak force is so weak, then who cares?
- The weak interaction is \*very\* important because it is the only one that can transform particles into other ones

![](_page_30_Figure_3.jpeg)

#### The Weak Interaction

 Without the weak interaction stars would not shine and there would be no light, heat, complex elements (oxygen, carbon, ... etc), planets.... life!

![](_page_31_Figure_2.jpeg)

• Also, it could be that neutrinos hold the key to some of the big answers, such as why is our universe made of matter.

#### How do you detect a neutrino?

- Reminder: of every100 000 000 000 000 000 000 solar neutrinos, only one in average interacts with a human being
- So detecting them is not impossible.... just very hard!
  - What should we do? Place a lot of mass in front of the neutrinos to increase the changes that at least a few will interact

![](_page_32_Picture_4.jpeg)

With a detector the size of one human being, 1 neutrino interacts

![](_page_32_Picture_6.jpeg)

With a detector the size of 15 human beings, ~15 neutrinos will interact

#### **Discovery of Neutrinos**

- The other thing you do: expose your detector to the largest possible flux of neutrinos
- This is why neutrinos were not discovered until 1956 by Frederick Reines and Clyde Cowan

![](_page_33_Picture_3.jpeg)

![](_page_33_Picture_4.jpeg)

#### Neutrino Detectors

- That's why neutrino detectors typically have to be HUGE
- Example: Super Kamiokande in Japan
  - Cylindrical tank 41m high and 39m in diameter
  - 50,000 tons of ultra pure water

![](_page_34_Figure_5.jpeg)

![](_page_34_Picture_6.jpeg)

## Super Kamiokande

The detector's interior is covered by approximately 11,146 photomultiplier tubes (extremely sensitive light sensors) visible in

the detector

![](_page_35_Picture_3.jpeg)

Each tube costs a few thousand dollars

![](_page_35_Picture_5.jpeg)

![](_page_36_Picture_0.jpeg)

#### Super Kamiokande

- Among the many things you can do with this detector: look at the sun at night
  - How? Only with neutrinos

![](_page_37_Picture_3.jpeg)

#### Super Kamiokande

• The sun at night as seen with neutrinos:

![](_page_38_Picture_2.jpeg)

Of course, we can do much more than just see the Sun: we can study it!

#### Kamiokande

- Neutrinos from other cosmic sources have also been detected:
  - On February 23 1987 a supernova exploded in the Large Magellanic Cloud

![](_page_39_Figure_3.jpeg)

- Kamiokande (Super Kamiokande's predecessor) and IMB
  detected 12 and 8 antineutrinos from this supernova, respectively
- If a supernova occurs now with our modern detectors, thousands of events will be detected

## **Reactor Neutrino Experiments**

- There are other types of neutrino experiments
- You can for example go near nuclear reactors

![](_page_40_Figure_3.jpeg)

![](_page_40_Figure_4.jpeg)

#### Example: Daya Bay

![](_page_41_Picture_1.jpeg)

#### Daya Bay

• The detectors are so sensitive they had to be assembled in a clean environment

![](_page_42_Picture_2.jpeg)

A few grams of sweat or dust could ruin everything!

![](_page_42_Picture_4.jpeg)

#### Accelerator Neutrino Experiments

• You can also create a beam of neutrinos and send it through the Earth

![](_page_43_Figure_2.jpeg)

- The beam is produced in Fermilab, near Chicago
- Neutrinos travel for 1,300 km under the Earth until they reach detectors with mass of 40,000 tons in an underground mine

## DUNE

#### One of DUNE's predecessors (MINOS)

DUNE is currently under construction:

![](_page_44_Picture_3.jpeg)

![](_page_44_Picture_4.jpeg)

- Here at UCI we are responsible for parts of the near detector and the purity monitors, among other areas (reconstruction and analysis)
- Start of operations: end of this decade
- The million dollar question is... why?

#### **Neutrino Oscillation**

• It turns out neutrinos have another mysterious property:

You don't have any friends. Nobody likes *you*.

![](_page_45_Picture_3.jpeg)

Not listening. Not listening.

#### **Neutrino Oscillation**

• Neutrinos change flavor as they travel!

![](_page_46_Picture_2.jpeg)

- How come?
  - It's a consequence of quantum mechanics, where particles behave like waves that interfere one with another
  - When created, neutrinos are not in a definite state of mass, but are instead a superposition of different mass states
  - Each mass state propagates slightly differently (in other words, each state is a wave with a slightly different frequency)

• Analogy: two musical notes that are slightly different (like when you tune a guitar)

![](_page_47_Figure_1.jpeg)

 Same thing with the probability of observing a certain type of neutrino (probability increases, decreases, increases, decreases... etc.)

#### Who cares about neutrino oscillation?

- Our best theory of particles and their interactions (the Standard Model) can explain almost every single observation done so far
  - This is both good and bad, because we know this theory is incomplete (does not have gravity, dark matter... etc)
  - Neutrino oscillation is one of those phenomena not explained by the Standard Model

#### Neutrinos can show us how to expand the theory!

- To incorporate neutrino oscillations, it is necessary to incorporate at least three new parameters to the actual theory, called mixing angles (θ<sub>12</sub>,θ<sub>23</sub>,θ<sub>13</sub>). It is also necessary to give neutrinos mass.
  - Let me finish by showing you one of our latest results

#### Latest Results from Daya Bay

- Latest results from Daya Bay
- We use more than 6
  million antineutrino
  interactions ——
- Can measure one of the mixing angles and one of the mass differences to < 3%.</li>

 $\sin^2 2\theta_{13} = 0.0853^{+0.0024}_{-0.0024}$  $\Delta m_{32}^2 = + \left(2.454^{+0.057}_{-0.057}\right) \times 10^{-3} \text{ eV}^2$ 

![](_page_49_Figure_5.jpeg)

#### Outlook

- There are still many open questions in neutrino physics:
  - Do neutrinos violate the chargeparity (CP) symmetry?
  - Are there more than three neutrinos?
  - How are neutrino masses ordered?

![](_page_50_Figure_5.jpeg)

- These questions have important consequences in physics & cosmology
  - Our group here at UCI is working hard to answer these and other questions

![](_page_50_Picture_8.jpeg)

- The Standard Model is our current best theory of particles and their interactions
  - It is extremely successful, perhaps too much so
- There are at least two ways of doing particle physics these days:
  - With collisions at colliders
  - With dedicated neutrino experiments (using accelerators, nuclear reactors, or natural sources)
- There are many open questions in our field
  - Neutrinos are one of the most promising ways of addressing some of these questions and of understanding the sources that produce them

## Questions?

![](_page_52_Picture_1.jpeg)

# Backup

#### The Higgs Mechanism

![](_page_54_Picture_1.jpeg)

### The Higgs Mechanism

![](_page_55_Picture_1.jpeg)

#### The Higgs Mechanism

![](_page_56_Picture_1.jpeg)

https://www.youtube.com/watch?v=joTKd5j3mzk