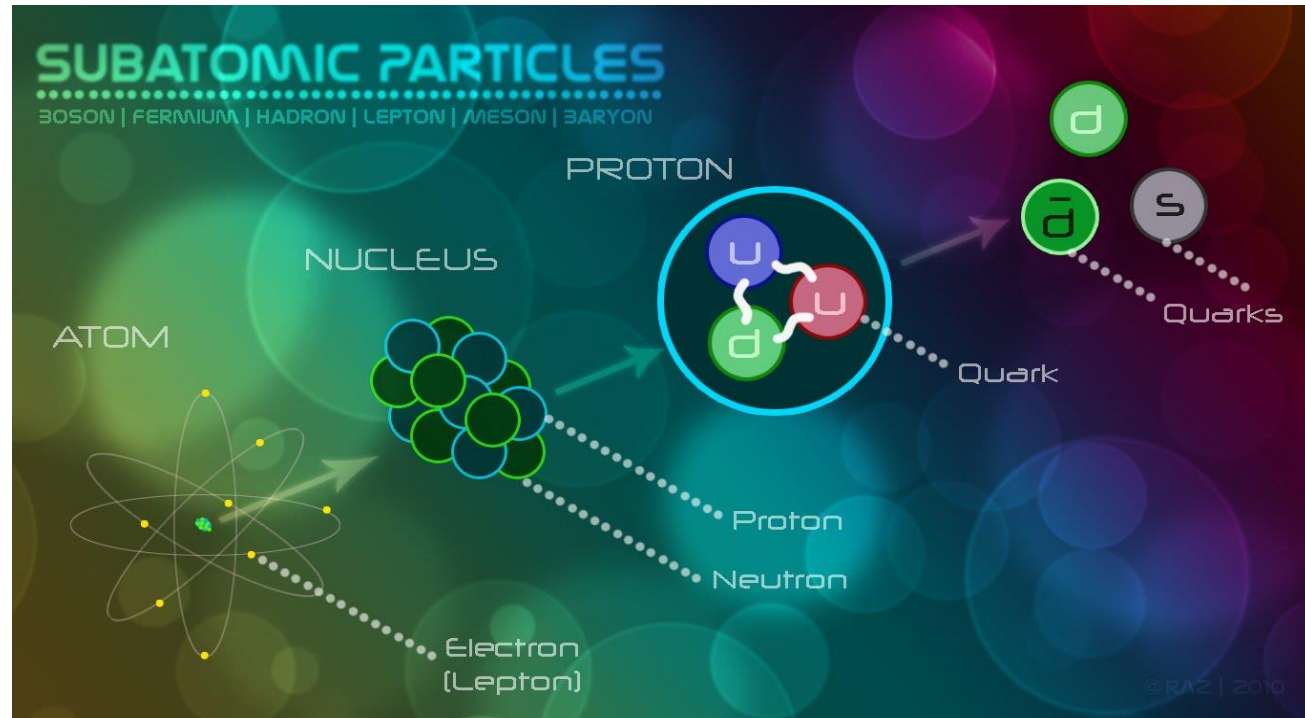


# The Standard Model of Particle Physics and Neutrino Physics

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# What is Particle Physics?

- Particle Physics is the study of the (subatomic) particles that constitute matter and radiation and their interactions.
- Particle Physics is the study of the really, truly, very, super-small stuff!



# These are not “classical particles”

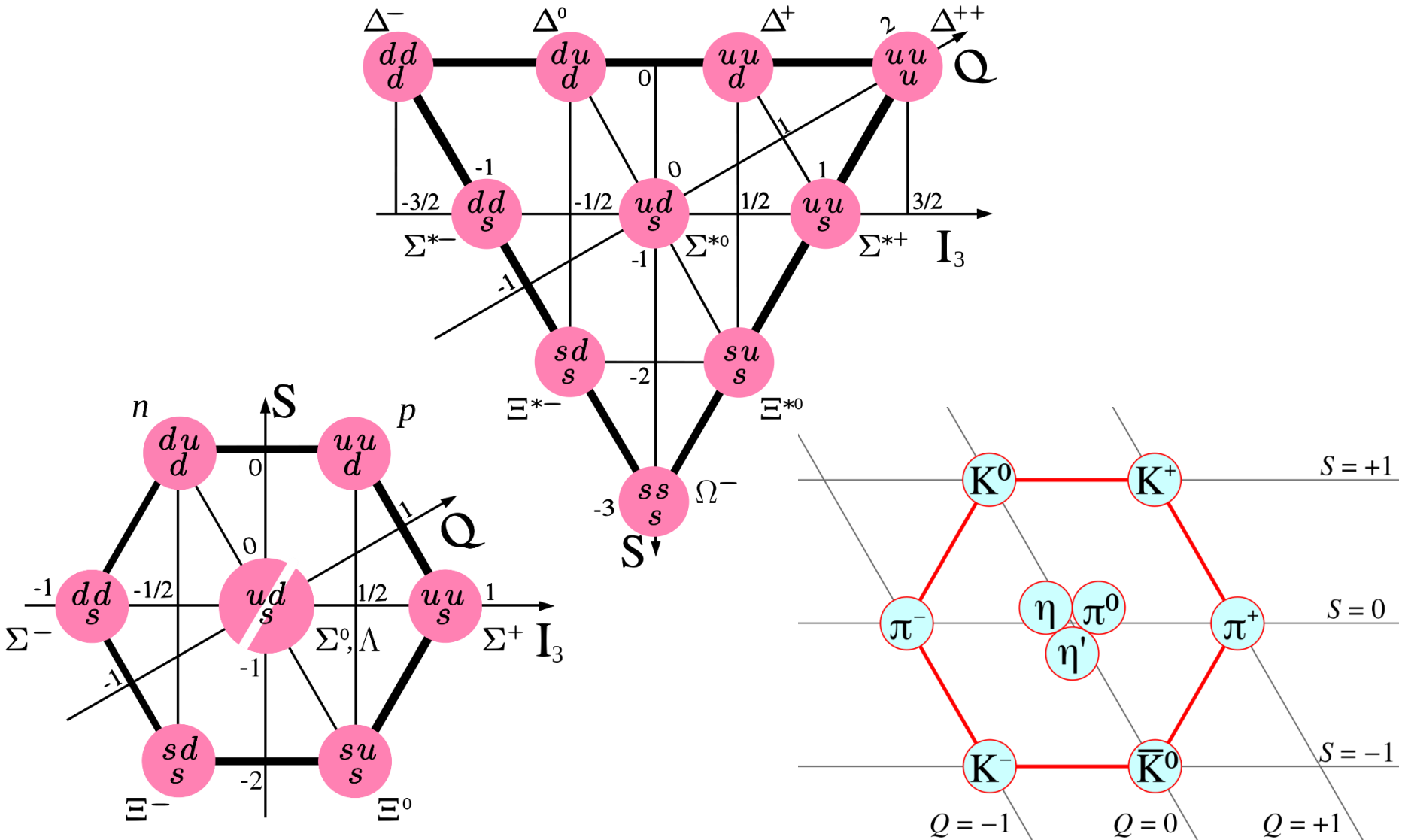
These particles obey the Principles of Special Relativity.

- The laws of Physics are the same in all inertial reference frames.
- The speed of light is the same in all inertial frames.

These particles obey the laws of Quantum Mechanics.

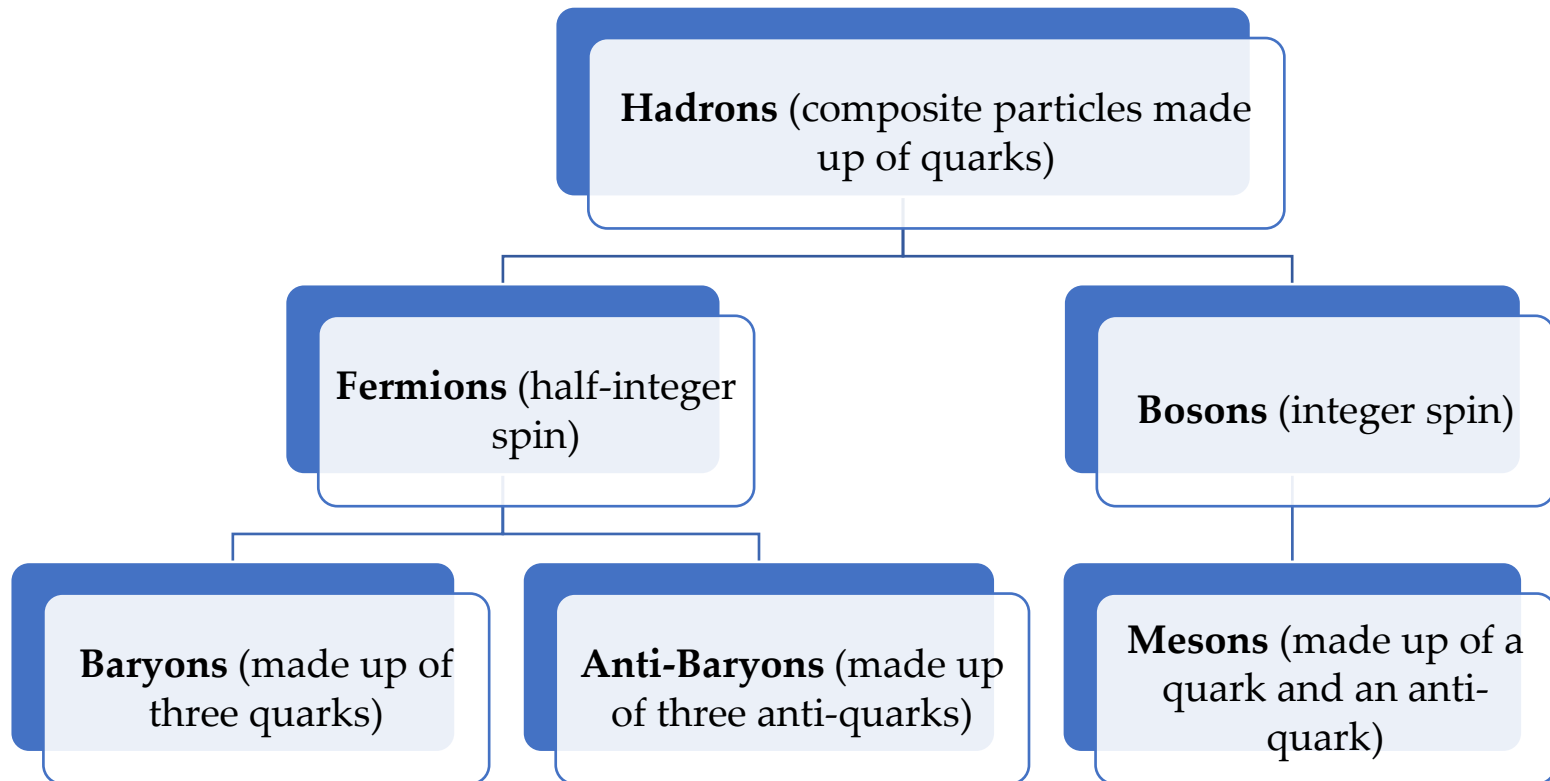
- The Heisenberg Uncertainty Principle.
- Probabilistic nature, subjective properties.
- Described by a wavefunction.

# What particles are there?

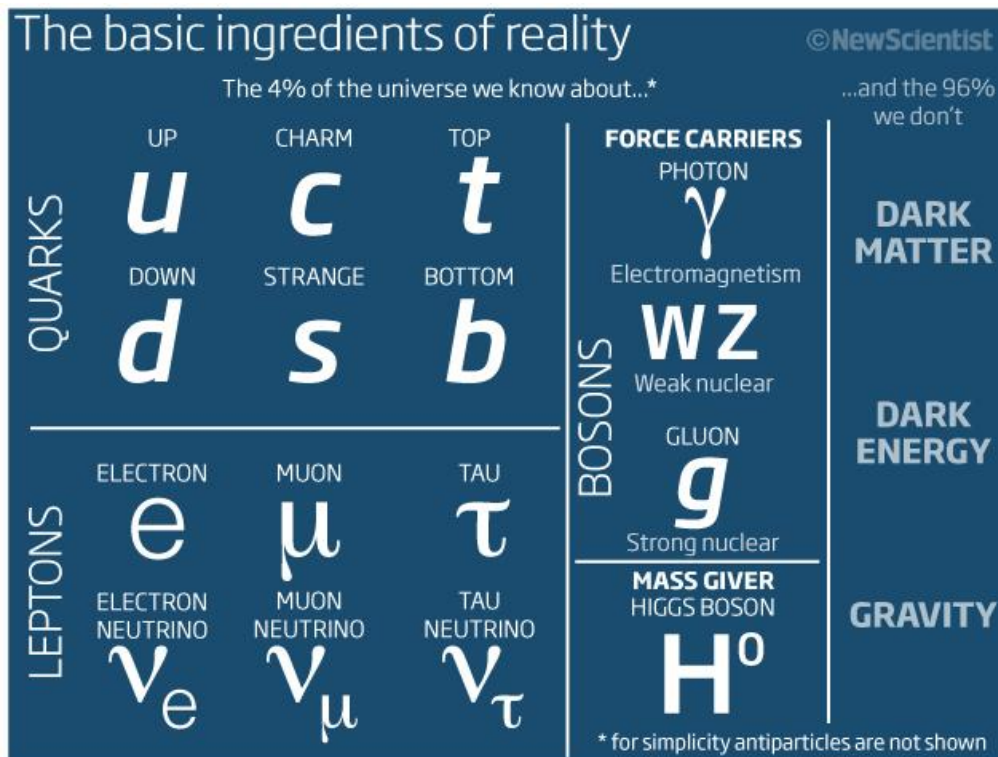


# How do we classify these particles?

- The theories of Special Relativity and Quantum Mechanics allow us to classify subatomic particles as fermions or bosons according to their spin.
- Particles made up of quarks are called hadrons and can be classified as baryons, anti-baryons and mesons.



# The Standard Model (SM)



- This theory allows us to classify all known subatomic particles in terms of a few **elementary particles**.
- Quarks and leptons (and their anti-particles) are fermions that make up all known matter.
- The photon, gluon, W and Z are bosons that “carry” the forces.
- The higgs boson gives (most?) particles their mass.

# What else does the SM do?

- The SM does more than classify particles, it is a mathematical theory that combines Special Relativity and Quantum Mechanics (our two best theories!) to explain nature at its most fundamental level.

$$\begin{aligned}
 \mathcal{L}_{SM} = & \underbrace{\frac{1}{4} W_{\mu\nu} \cdot W^{\mu\nu} - \frac{1}{4} B_{\mu\nu} \cdot B^{\mu\nu} - \frac{1}{4} G_{\mu\nu} \cdot G^{\mu\nu}}_{\text{Kinetic energies and self-interactions of the gauge bosons}} \\
 & + \underbrace{\bar{L}\gamma^\mu \left( i\partial_\mu - \frac{1}{2} g\tau \cdot W_\mu - \frac{1}{2} g'Y B_\mu \right) L + \bar{R}\gamma^\mu \left( i\partial_\mu - \frac{1}{2} g'Y B_\mu \right) R}_{\text{Kinetic energies and electroweak interactions of fermions}} \\
 & + \underbrace{\frac{1}{2} \left| \left( i\partial_\mu - \frac{1}{2} g\tau \cdot W_\mu - \frac{1}{2} g'Y B_\mu \right) \phi \right|^2 - V(\phi)}_{\text{W}^\pm, Z, \gamma \text{ and Higgs masses and couplings}} \\
 & + \underbrace{g''(\bar{q}\gamma^\mu T_a q)G_\mu^a}_{\text{interactions between quarks and gluons}} + \underbrace{(G_1 \bar{L}\phi R + G_1 \bar{L}\phi_c R + h.c.)}_{\text{fermion masses and couplings to Higgs}}
 \end{aligned}$$

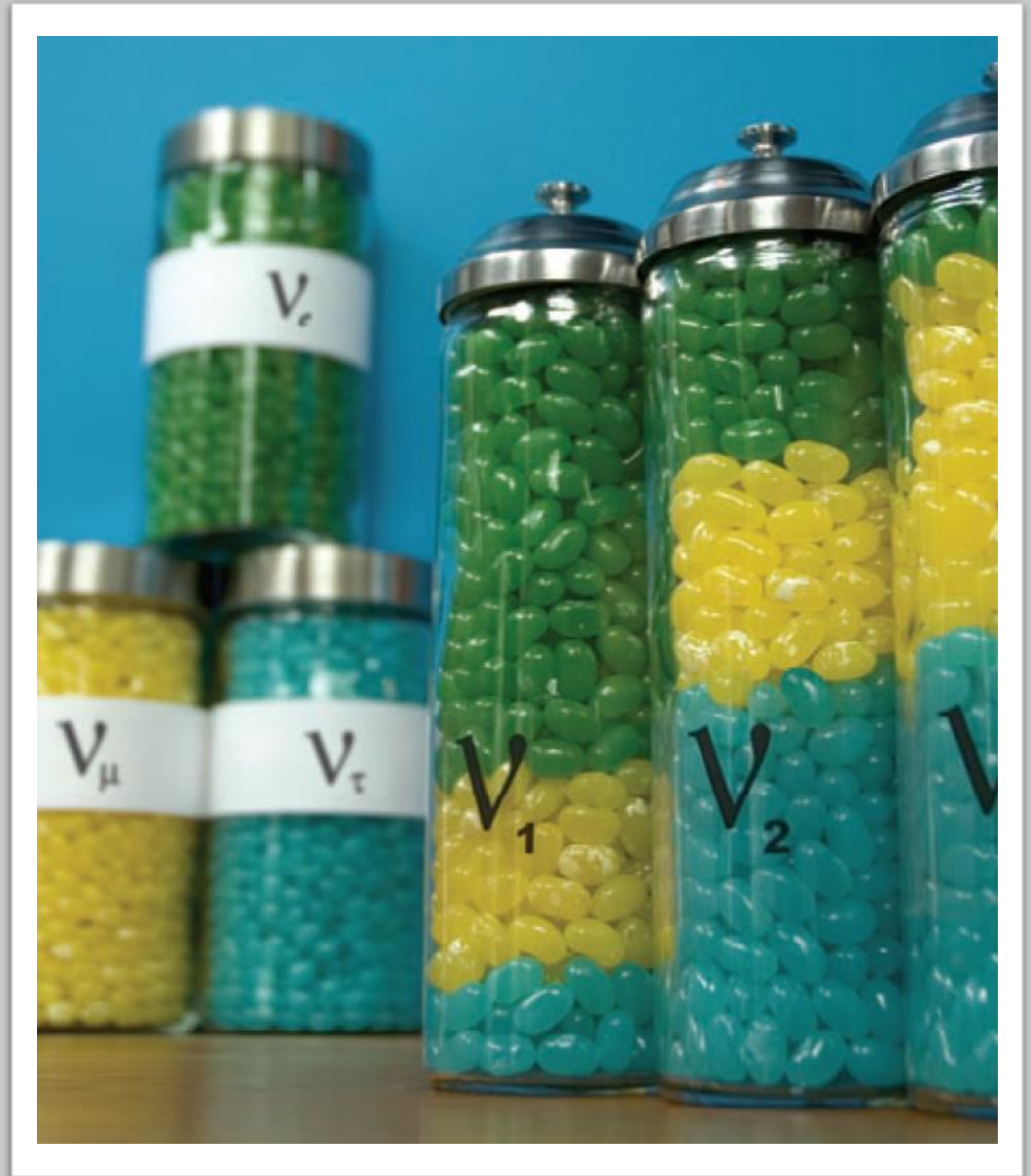
Some things we  
know about  
neutrinos

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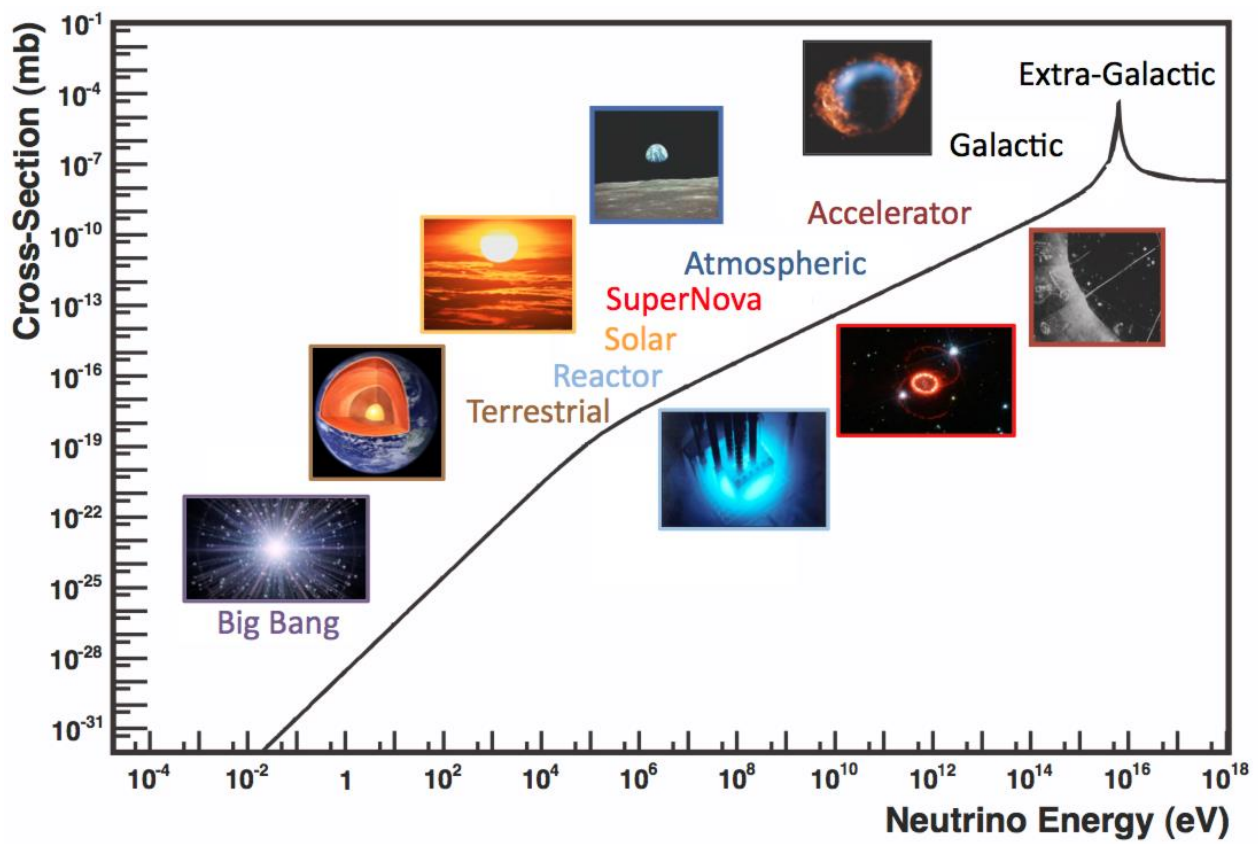


# There are three types of neutrinos

- Electron neutrinos
- Muon neutrinos
- Tau neutrinos



# They come from different sources



# They are everywhere

Every  $\text{cm}^2$  of Earth surface is crossed every second by more than 10 billion ( $10^{10}$ ) neutrinos produced in the Sun

Within your body at any instant: roughly 30 million neutrinos from the Big Bang

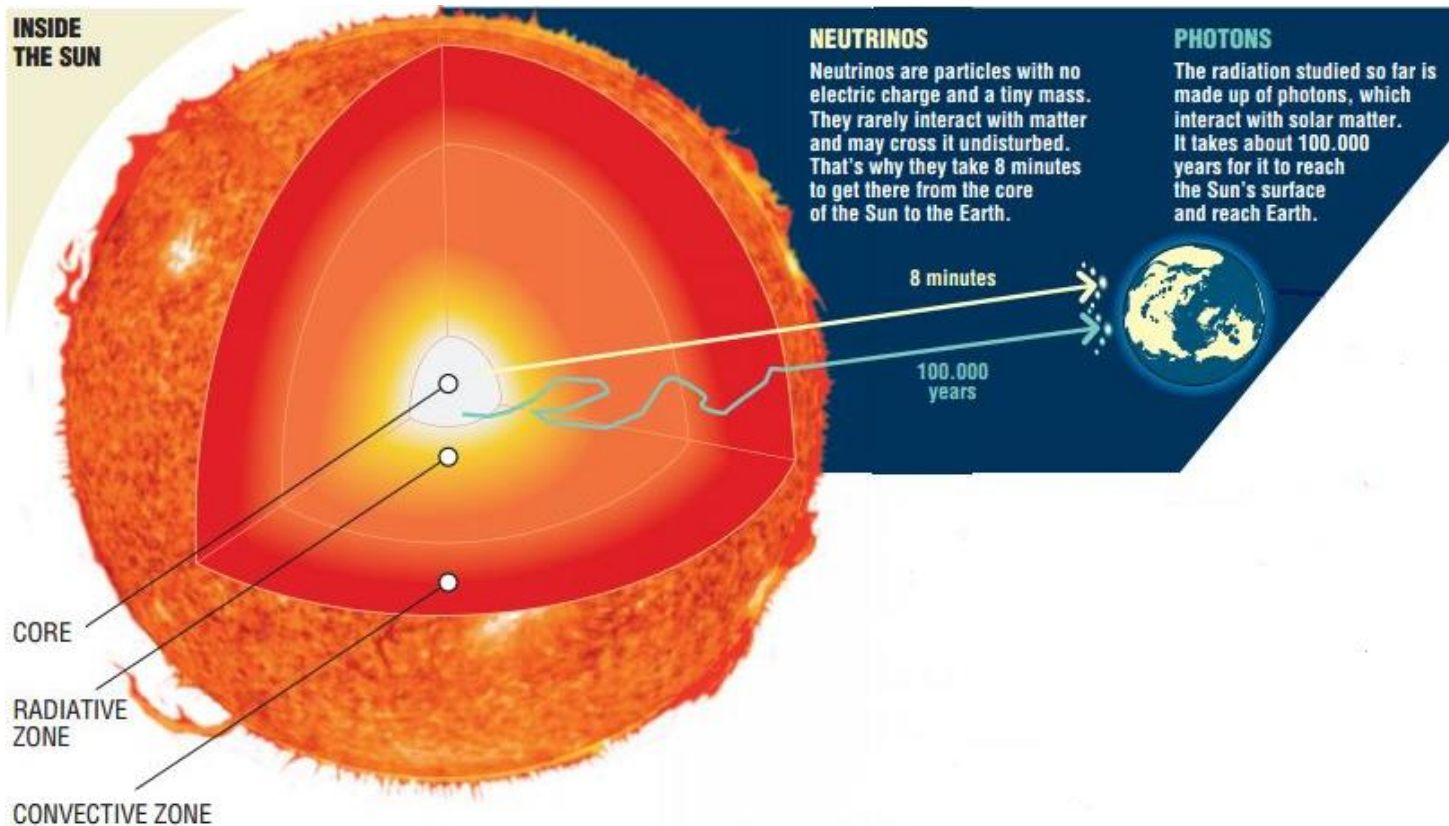


$10^{14}$  neutrinos per second from Sun are zipping through you

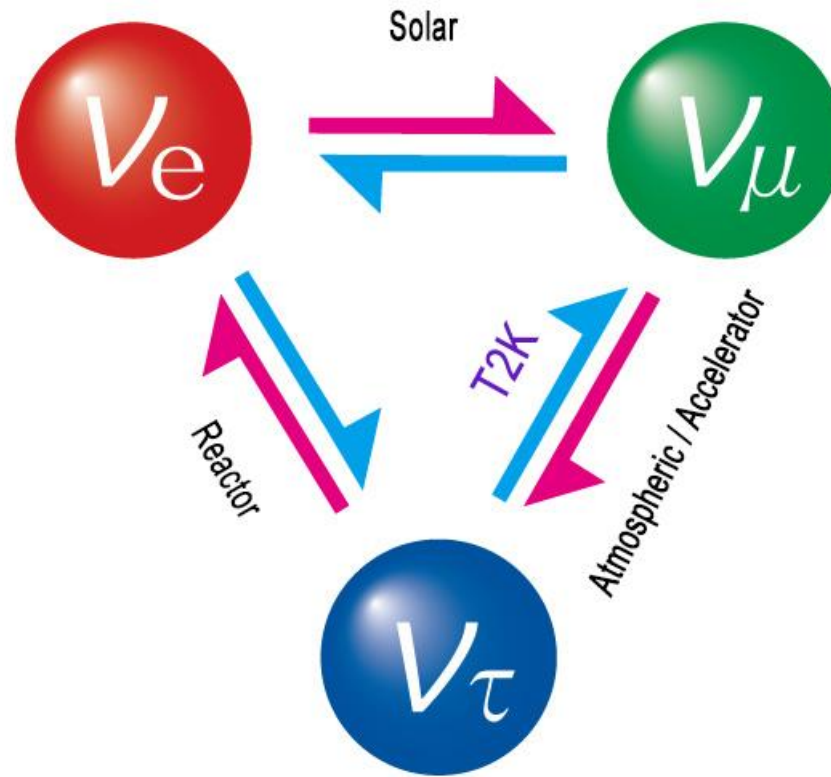
*No worries!  
Neutrinos do not harm us.  
Our bodies are transparent to neutrinos*

# neutrinos are very weakly interacting

- No charge
- No color
- Almost no mass



# Neutrinos have (very small) mass...and oscillate!



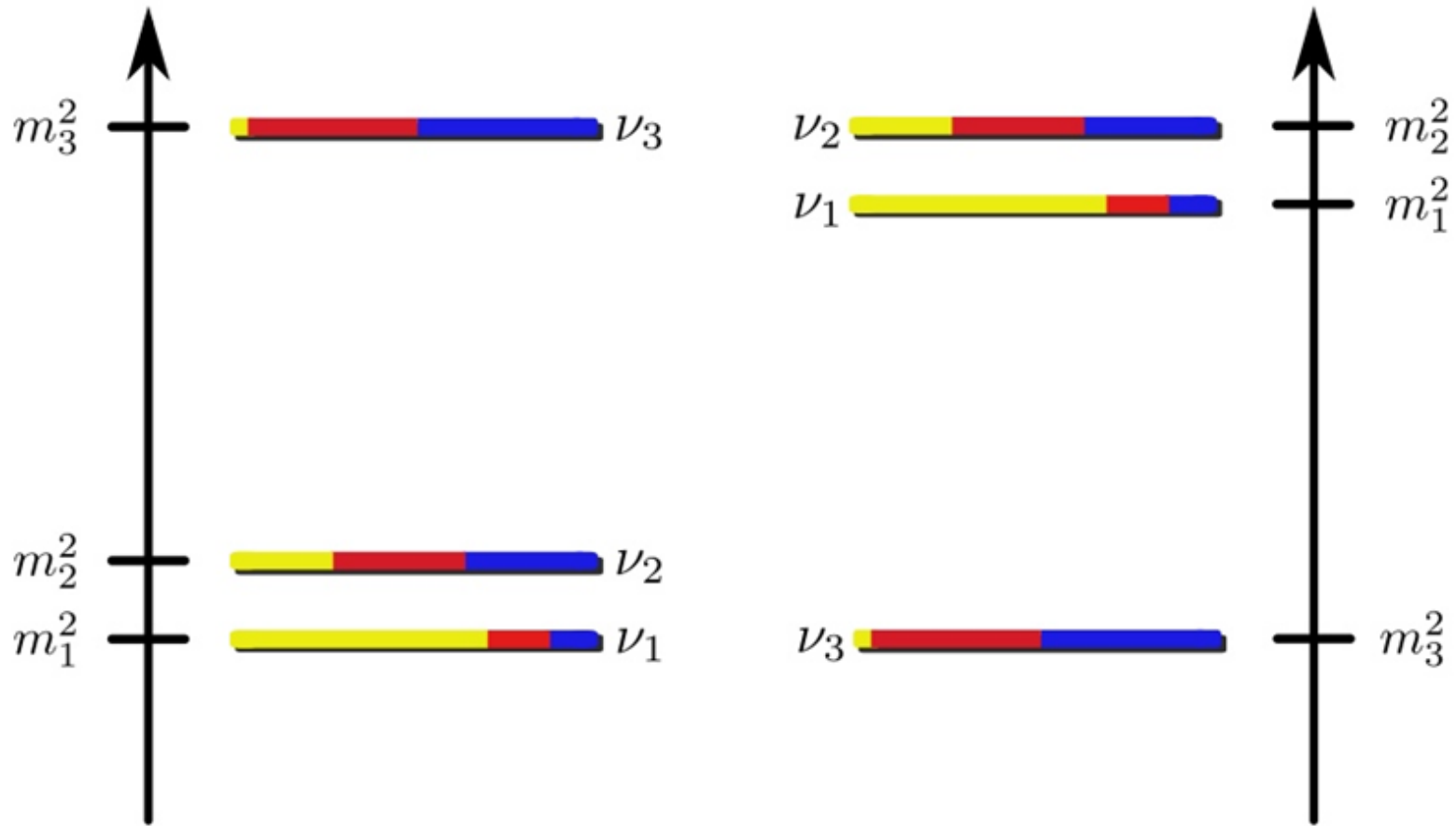
Neutrino oscillation between three generations

(some of)  
what we don't know

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# Mass Hierarchy



# Missing parameters





$$\begin{aligned}
 U &= \begin{bmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{bmatrix} \\
 &= \begin{bmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{bmatrix} \begin{bmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{bmatrix} \begin{bmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} e^{i\alpha_1/2} & 0 & 0 \\ 0 & e^{i\alpha_2/2} & 0 \\ 0 & 0 & 1 \end{bmatrix} \\
 &= \begin{bmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{bmatrix} \begin{bmatrix} e^{i\alpha_1/2} & 0 & 0 \\ 0 & e^{i\alpha_2/2} & 0 \\ 0 & 0 & 1 \end{bmatrix}
 \end{aligned}$$



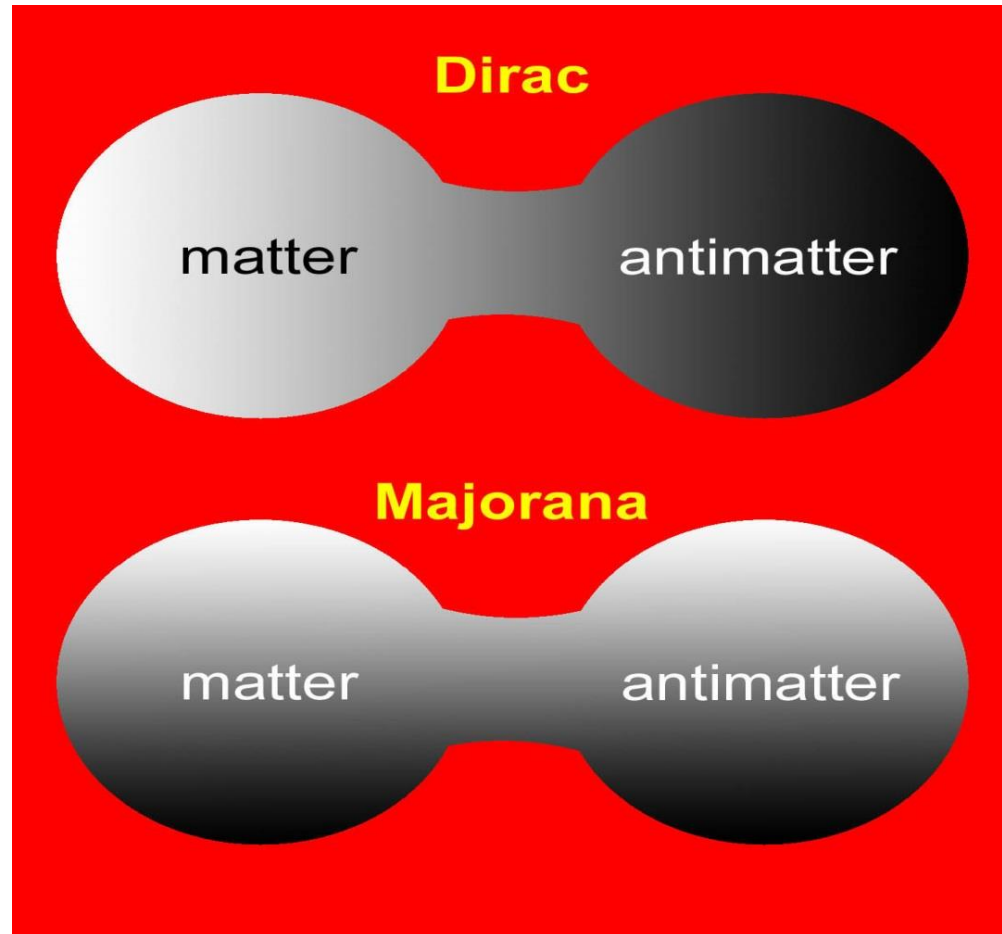
# Are there more neutrinos?

Desperately seeking sterile

The three known types of neutrino might be "balanced out" by a bashful fourth type

ELECTRON NEUTRINO	MUON NEUTRINO	TAU NEUTRINO	STERILE NEUTRINO
			
MASS	< 1 electronvolt		>1 electronvolt
FORCES THEY RESPOND TO	Weak force Gravity		Gravity
DIRECTION OF SPIN	All three "left handed"		"Right handed"

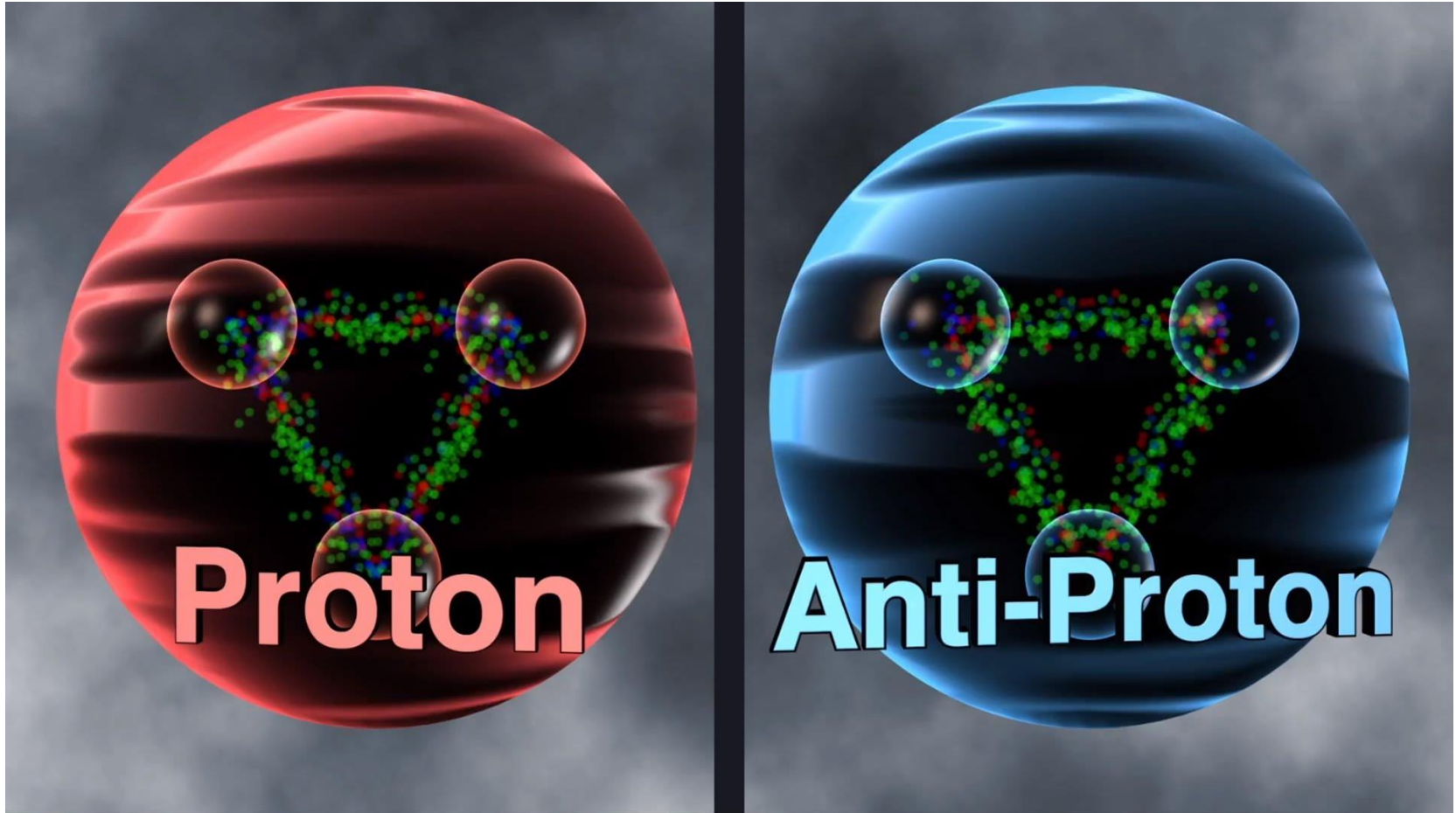
# What kind of mass do they have?



Why are neutrinos  
important?

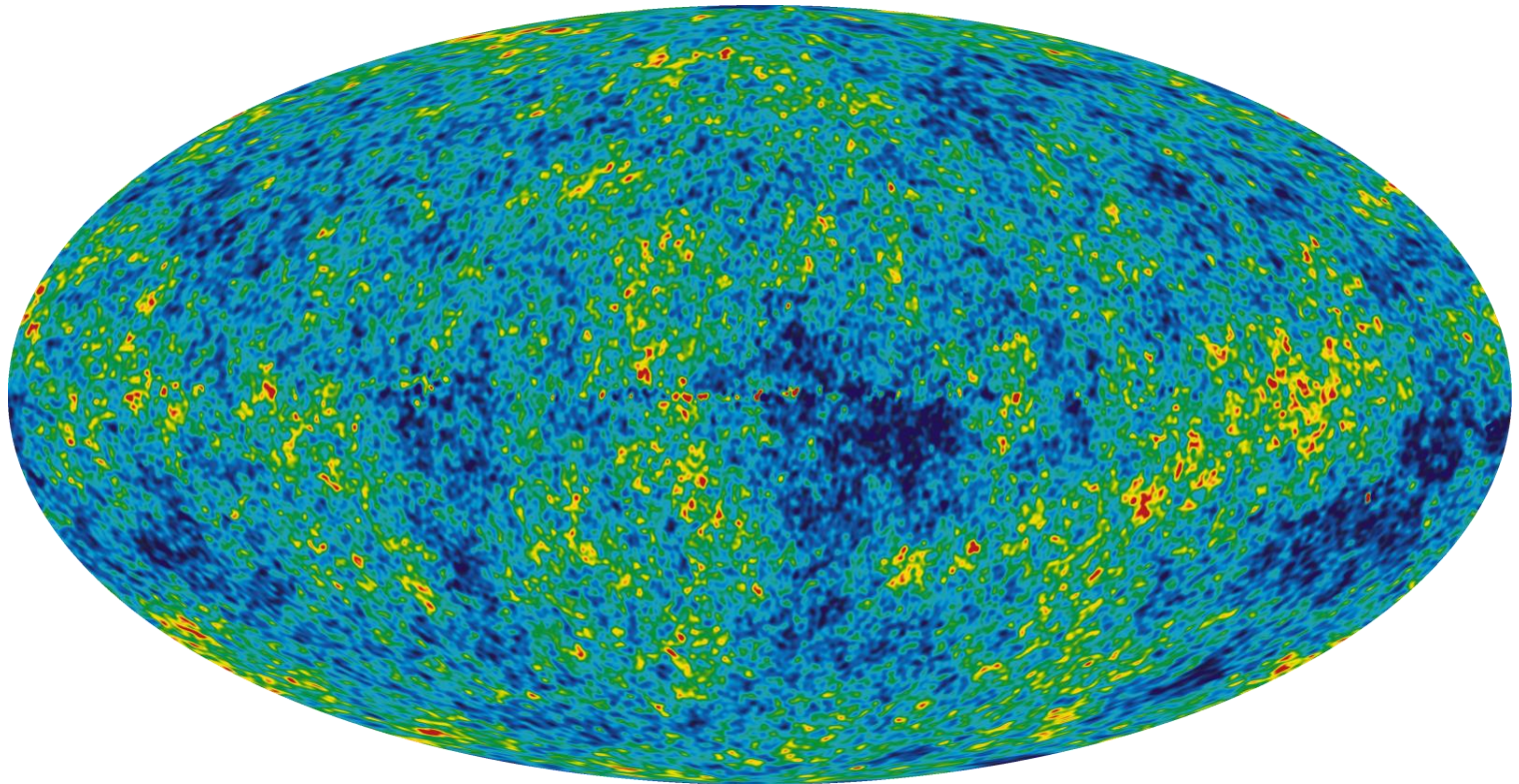
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Neutrinos are probably responsible for the origin of baryon asymmetry in the universe.



Neutrinos are important for

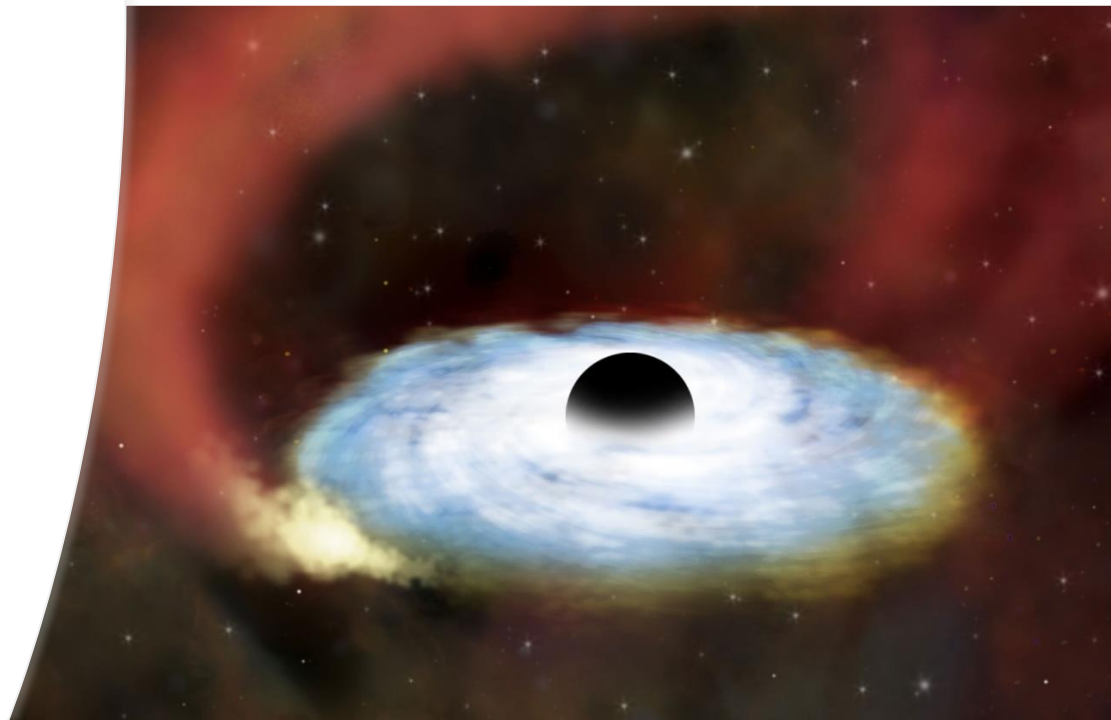
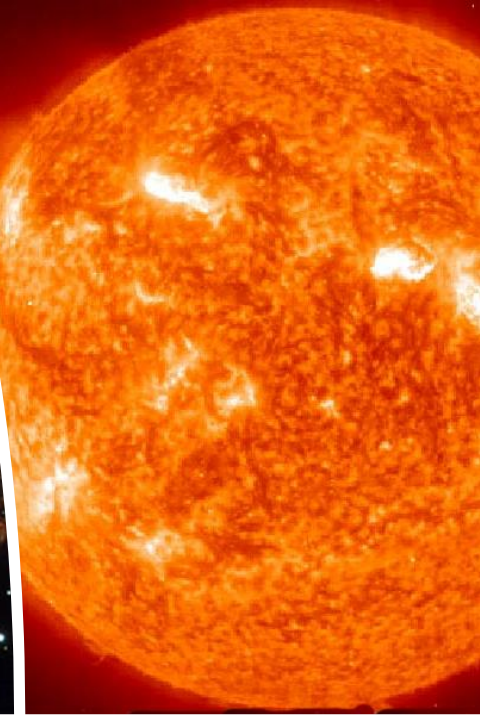
- understanding the spectrum of the CMB
- making sense of the LSS in the universe
- studying the delicate chemical equilibriums determining the light element abundances during BBN





Neutrinos could be a powerful probe of

- the dynamics of stars
- supernova collapse
- accretion disks encircling supermassive black holes.



Thank you for your  
attention